Black Hole Formation from High Energy Scattering in AdS/CFT

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Outline

- I. How to create a Black Hole in AdS.
 - o The flat space picture.
 - o Graviton scattering in AdS and the 1/2 BPS geometries (LLM).
- II. The gauge theory (initial) states.
- III. The strong coupling description (a proposal)
 - o Matrix quantum mechanics

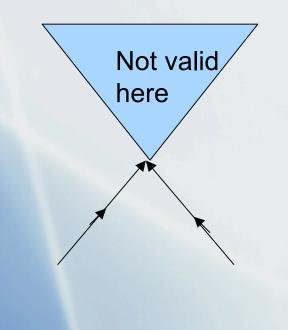
Main reference: arXiv:0709.3503

1. Consider two massless particles on a head-on collision. o To form a classical black hole we need $E_{cm} >> E_{planck}$.

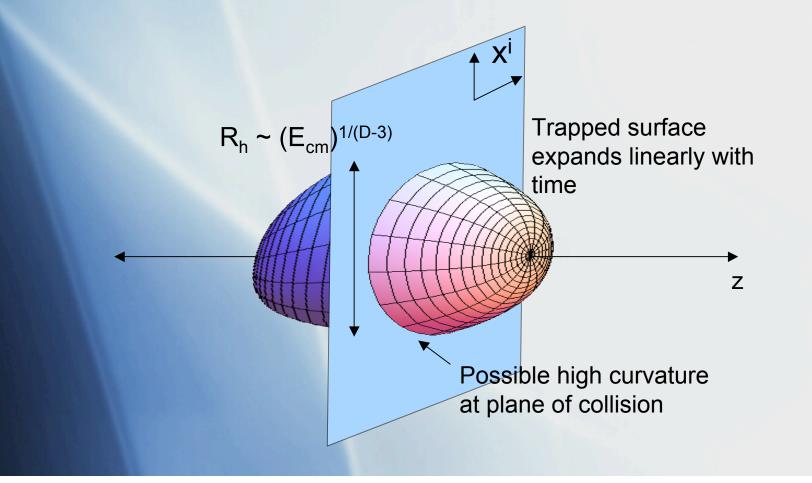
2. To study such collision (classicaly) one can use the Aichelburg-Sexl metric:

$$u = t - z$$
, $v = t + z$.

- 1. This metric can be obtained by boosting the Schwarzchild metric and taking the mass to zero.
- 2. One can superpose two shock waves and form a solution to Einstein's equations outside the future light cone of the collision.



1. This superposition leads to the formation of a Marginally Trapped Surface (Giddings, Eardley hep-qc/0201034)



1. Danger with high curvatures: For a single shock,

$$R_{uivj} = -rac{1}{2}\delta(u)\partial_i\partial_j\Phi(ec{x})$$

but $R_{\mu\nu\rho\sigma}R^{\mu\nu\rho\sigma}$ is still finite.

2. For two shocks we have divergence of the form

$$R_{\mu\nu\rho\sigma}R^{\mu\nu\rho\sigma} \sim \frac{E^2}{|\vec{x}|^{2D-4}}\delta(u)\delta(v)$$

3. Need to regularize delta-function (by creating wave-packet)

$$\delta(u) \to \delta_{\lambda}(u)$$

1. For low curvatures we need

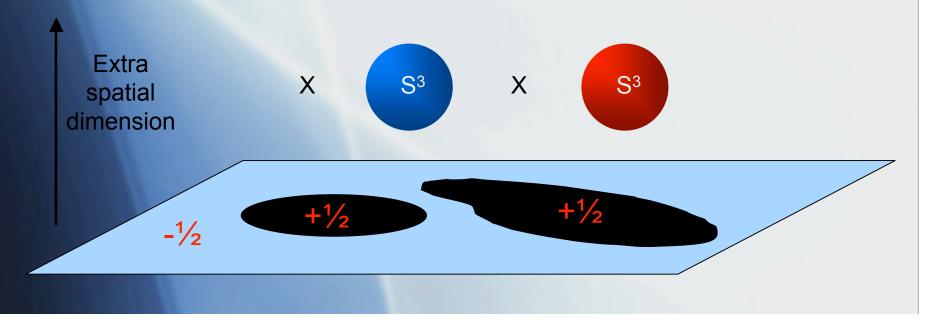
 $(R_{\mu
u
ho\sigma})^2 \ll 1/l_s^4 \quad \left| ec{r}_h \right| \gg \lambda \quad \left| ec{r}_h \right| \gg l_s$

2. This gives

•
$$g^{-1/2}E^{-1/7} \ll \lambda \ll E^{1/7}$$
,
• $E^{1/7} \gg g^{-1/4}$.

Conventions. In this talk we will set the ten dimensional Planck's constant $l_p = 1$. In these units, the AdS radius is related to the rank of the SYM gauge group N by $R = (4\pi N)^{1/4}$. The string length is given by $l_s = \sqrt{\alpha'} = g^{-1/4}$, where g is the closed string coupling also related to the SYM coupling by $4\pi g = g_{YM}^2$. We will also introduce the parameter $\hbar = 1/N$.

- Here we want to explain how to get the smooth graviton wave-packets from the 1/2 BPS solutions of type IIB SUGRA.
 - o N units of RR five-form flux
 - o All solutions asymptotically $AdS_5 \times S^5$
 - Classified by a single function that takes values ± 1/2 on a twodimensional plane. (Lin, Lunin, Maldacena, 2004)



1. Metric:

V

$$ds^{2} = -h^{-2}(Dt)^{2} + h^{2}(dy^{2} + dzd\bar{z}) + ye^{-G}d\Omega_{3}^{2}$$
$$+ye^{G}d\tilde{\Omega}_{3}^{2},$$
$$h^{-2} = 2y\cosh G,$$
$$f = \frac{1}{2} \tanh G,$$
$$f(z, \bar{z}, y) = -\frac{y^{2}}{2} \int d^{2}z' \frac{\rho(z', \bar{z}')}{(|z - z'|^{2} + y^{2})^{2}}.$$
$$Dt = dt + V = dt + \frac{1}{2}i\bar{V}dz - \frac{1}{2}iVd\bar{z}$$
$$V(z, \bar{z}, y) = \frac{1}{2} \int d^{2}z' \frac{\rho(z', \bar{z}')(z - z')}{(|z - z'|^{2} + y^{2})^{2}}.$$

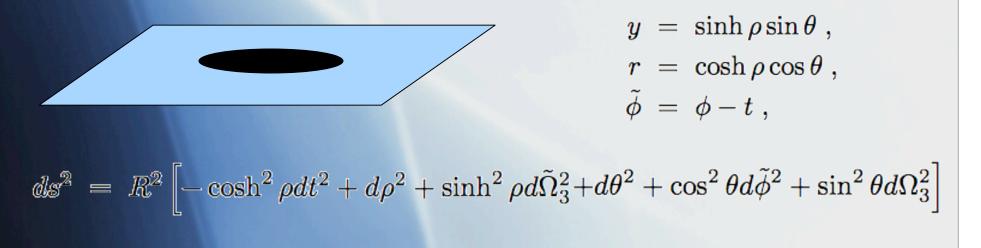
1. The area of the droplet is constrained by

$$\int_{\mathcal{D}} \frac{d^2 z}{\pi} = 1$$

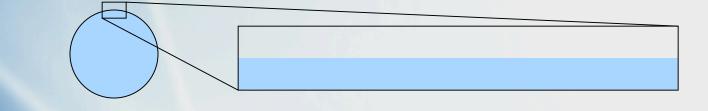
2. The energy of the solution is

$$E=J=rac{1}{\hbar^2\pi}\int_{\mathcal{D}}d^2z|z|^2-rac{1}{2\hbar^2}\,.$$

3. Example: $AdS_5 \times S^5 = unit disk$



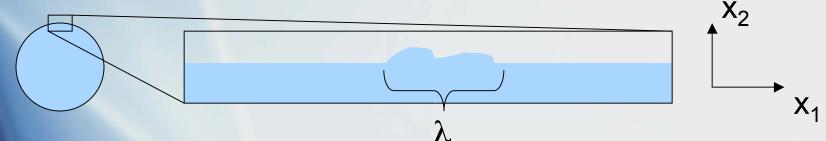
1. To get flat space we zoom in into the edge of the droplet (the equator of the S^5 and the origin of AdS_5).



2. Do this by rescaling coordinates and taking R-> ∞ , $z \to \left(1 + \frac{x_2}{R^2}\right) \exp\left(\frac{\pi}{2} - \frac{x_1}{R}\right), \ y \to \frac{1}{R^2}y, \ t \to \frac{1}{R}t$ We get, ($y = r_1r_2, x_2 = (r_1^2 - r_2^2)/2$) Note that to space limit fixed, so the large N limit fixed is so the large N limit for the space limit fixed is so the large N limit for the space limit fixed is so the large N limit for the space limit for the space limit fixed is space limit fixed is so the large N limit for the space N l

Note that to get the flat space limit, we keep I_s fixed, so this is taking large N limit with g fixed but small. (not 't Hooft limit)

1. To create a graviton, we make a small ripple on the edge of the droplet: (note the two geodesics $x_1 = \text{const.}$, $x_1 = 2 \text{ t} + \text{const.}$)



2. Parametrize deformation of edge as,

 $r_{
m boundary} = 1 + \delta r(\phi)$

3. The energy is,

$$E \approx \frac{1}{\pi \hbar^2} \int d\phi \left[\delta r(\phi) \right]^2$$

1. To keep the energy fixed as we take R to infinity, the perturbation must take form

$$\delta r(\phi) \;=\; rac{4\pi^{3/2}\sqrt{E}}{R^3\sqrt{\lambda}}g(x_1/\lambda) \;.$$

2. Without lost of generality, g(x) is zero outside the interval [-1/2, 1/2], and it's normalized as

$$\int_{-1/2}^{1/2} dx [g(x)]^2 = 1 \qquad \int_{-1/2}^{1/2} dx g(x) = -\frac{2\pi^{3/2}\sqrt{E}}{R^3\sqrt{\lambda}} \to 0$$
(Area conservation)

1. One can show that, after taking R to infinity, we get a regularized Aichelburg-Sexl metric in 10 dim.

$$ds^{2} = -dx_{1}(2dt - dx_{1}) + d\vec{r}^{2} + \frac{(4\pi)^{3}E}{|\vec{r}|^{6}}\delta_{\lambda}(x_{1})dx_{1}^{2}$$

where

$$\delta_\lambda(x) = \int_{-1/2}^{1/2} dx' [g(x')]^2 \delta(x-\lambda x')$$

2. We can put now an anti-1/2 BPS particle traveling in the opposite direction by replacing $x_1 \rightarrow 2t - x_1$.

Gauge Theory Interpretation

- 1. Now we want to find the gauge theory dual of the regularized Aichelburg-Sexl metric and learn how to set up the initial states for the graviton scattering.
- 2. 1/2 BPS states of SYM theory: zero modes of one complex scalar

SYM on R x S³

•Expand in spherical harmonics

$$Z(t,\Omega) = \sum_{A} Z_A(t) Y_A(\Omega)$$

•Effective tree level action for zero mode (A = 0)

$$S = \int dt \, \operatorname{Tr}\left(|\dot{Z}|^2 - |Z|^2
ight)$$

Gauge Theory Interpretation

1. Define the operators

$$A^\dagger = rac{1}{\sqrt{2}} \left(Z - i ar{\Pi}
ight) \;, \qquad ar{A}^\dagger = rac{1}{\sqrt{2}} \left(ar{Z} - i \Pi
ight)$$

with

$$\begin{split} [A_i^j,(A^\dagger)_k^l] &= \delta_i^l \delta_k^j \ , \quad [\bar{A}_i^j,(\bar{A}^\dagger)_k^l] = \delta_i^l \delta_k^j \ , \end{split}$$
 Note that $\Pi^\dagger = \bar{\Pi}.$

2. The Hamiltonian and R-charge operator are

 $H = \operatorname{Tr} \left(A^{\dagger}A + ar{A}^{\dagger}ar{A}
ight) \qquad J = \operatorname{Tr} \left(A^{\dagger}A - ar{A}^{\dagger}ar{A}
ight)$

3. Since [H,J] = 0, we can define

$$H' = H - J$$

Identify with global time in AdS

Gauge Theory Interpretation

1. The (anti) 1/2 BPS operators have the form,

 $|\psi_{1/2\mathrm{BPS}}
angle = \psi(A^{\dagger})|0
angle = \psi(ar{A}^{\dagger})|0
angle$

where $(H \pm J) |\psi\rangle = 0$

Relation to Graviton Scattering:

1. An initial state for a (head-on) scattering process will take the form:

$$|\psi
angle \propto e \; {
m Tr} \Omega_1(ar A^\dagger) e \; {
m Tr} \Omega_2(A^\dagger) |0
angle$$

1. To find the geometric interpretation of the 1/2 BPS states introduce a coherent state:

$$A_i^j |Z
angle \ = \ Z_i^j |Z
angle$$

2. The 1/2 BPS state becomes

$$\langle Z|\psi\rangle = e^{\operatorname{Tr}\Omega(Z)/\hbar}e^{-\operatorname{Tr}|Z|^2/(2\hbar)}$$

3. The normalization is given by a complex random-matrix integral

$$\langle \psi | \psi
angle ~=~ \int [d^2 Z] |\langle \psi | Z
angle |^2$$

1. We can now go to an eigenvalue basis (I have rescaled $Z \rightarrow Z/\sqrt{\hbar}$):

From / Jacobian

$$\langle \psi | \psi \rangle \propto \prod_{i=1}^{N} \int d^2 z_i e^{\sum_j W(z_j, \bar{z}_j)/\hbar + \sum_{i < j} \log |z_i - z_j|^2}$$

where

$$W(z, \overline{z}) = -|z|^2 + \Omega(z) + \overline{\Omega(z)}$$

2. Taking the large N limit ($\hbar \rightarrow 0$) we can use the saddle point approximation and replace sums by integrals over eigenvalue distributions

1. Saddle point equations:

 $\delta F[\rho] = 0$

where

$$\begin{split} F[\rho] \;\;=\;\; -\frac{1}{\hbar} \int d^2 z \rho(z) W(z,\bar{z}) \\ &\;\; -\frac{1}{2} \int d^2 z \int d^2 z' \rho(z) \rho(z') \log |z-z'|^2 \end{split}$$

 $ho = 1/(\hbar\pi)$

and subject to the constraint $\int d^2 z \rho = \hbar^{-1}$

2. This gives constant density domains ("droplets") of eigenvalues:

- 1. These are the 1/2 BPS geometries!
- 2. A precise dictionary for a single droplet was developed in (Vazquez hep-th/0612014)
 - o Write potential as,

$$\Omega(z) = \sum_{k>0} t_k z^k$$

then

$$t_k = \frac{1}{k} \oint_{\partial \mathcal{D}} \frac{dz}{2\pi i} \bar{z} z^{-k} \quad \longleftrightarrow \quad \mathcal{D}$$

3. The energy of the state coincides with SUGRA result:

$$E = rac{1}{\hbar} \langle \mathrm{Tr} |Z|^2
angle - rac{1}{2\hbar^2} pprox rac{1}{\hbar^2 \pi} \int_{\mathcal{D}} d^2 z |z|^2 - rac{1}{2\hbar^2}$$

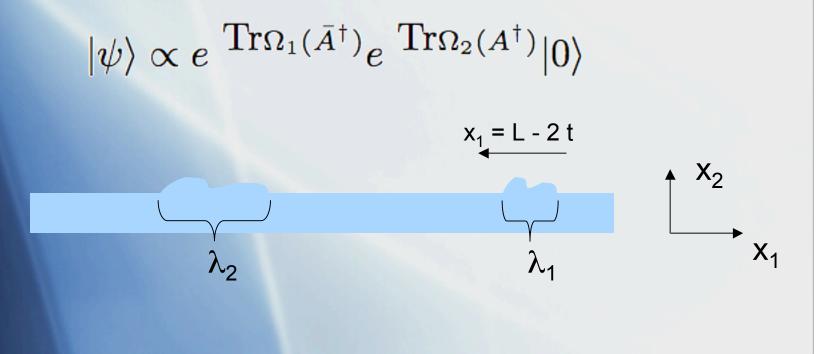
1. For a linearized perturbation around the circular droplet

$$t_k \approx \frac{1}{\pi k} \int_0^{2\pi} d\phi \delta r(\phi) e^{-ik\phi}$$

2. So we can now write a dual to the Aichelburg-Sexl geometry:

$$ert \psi
angle \ = \ \exp \left[-\sqrt{rac{E\lambda}{\pi}} \int_{-1/2}^{1/2} dx g(x) \ ext{Tr} \log \left(i e^{-ix\lambda/R} - \sqrt{\hbar} A^{\dagger}
ight)
ight] ert 0
angle \ ds^2 \ = \ -dx_1 (2dt - dx_1) + dec r^2 + rac{(4\pi)^3 E}{ec r ec ec ^6} \delta_\lambda(x_1) dx_1^2$$

- 1. The dual of a shock wave at $x_1 = L$ (at t = 0) can be obtained by shifting $x \rightarrow x + L$ and $A^{\dagger} \rightarrow \overline{A}^{\dagger}$ in the previous state.
- 2. So long as the initial waves a sufficiently separated in spacetime, one can write the initial state as a product of 1/2 and anti-1/2 BPS:



1. To ensure that the resulting collision leads to the formation of a classical black hole, we need, again

•
$$g^{-1/2}E^{-1/7} \ll \lambda \ll E^{1/7}$$
,
• $E^{1/7} \gg g^{-1/4}$.

2. This is very easy to satisfy...

Note that to get the flat space limit, we keep I_s fixed. This is taking the large N limit with g fixed but small. (not 't Hooft limit)

- 1. So far, we have seen that we can set up initial states that, according to the dual semiclassical gravity, will result in the formation of classical black holes in the bulk.
- 2. Now we want to know how could we describe this process at strong coupling.
- 3. Proposal: Matrix Quantum Mechanics
 - 1. The low energy dynamics of SYM theory seems to be described by a reduced model of matrix quantum mechanics: Berenstein hep-th/0403110, hep-th/0507203.
 - 2. The lowest energy states of SYM are the susy or BPS states (vacua).
 - 3. Using the operator state correspondence, one can classify vacua according to which operator acquires a vev in flat space SYM.
 - 4. On the BPS states, it turns out that the only operators that acquire a v.e.v. are scalars

 $\mathcal{O} \sim \operatorname{Tr}(XYZ\cdots)$

1. Moreover, the F-term conditions are zero in the chiral ring:

 $\langle \partial_{\Phi^i} W(\Phi) \mathcal{O}_1 \mathcal{O}_2 \cdots
angle = 0$

where the superpotential is

 $W = \operatorname{Tr}(X[Y, Z])$

2. This means that all expectation values containing commutators are zero in the chiral ring

$$[X,Y] = 0$$
, $[X,Z] = 0$, $[Y,Z] = 0$

1. In particular, operators that differ by a commutator have same vev:

 $\langle \operatorname{Tr}(XYZ\cdots)\rangle = \langle \operatorname{Tr}(XZY\cdots)\rangle$

- Since the scalar operators are dual to the zero mode of the same scalar in S³ (for SYM on R x S³), this suggests that the dynamics of the BPS sector can be described by a reduced model including only these fields.
- 2. Effective tree-level Hamiltonian:

$$H_{eff} = \operatorname{Tr}\left(\sum_{a=1}^{6} \frac{1}{2} (D_t X^a)^2 + \frac{1}{2} (X^a)^2 + \sum_{a,b=1}^{6} \frac{g_{YM}^2}{8\pi^2} [X^a, X^b] [X^a, X^b] \right)$$

3. We now want to impose the F-term conditions:

 $[X^a, X^b] = 0$

- 1. Another way of seeing the meaning of F-term condition is that we are taking low energy limit:
 - From AdS/CFT and general gauge theory considerations we know that the "size" of the ground state is order N²

$$rac{1}{N}\sum_{a=1}^6 \langle 0 | \ {
m Tr}(X^a X^a) | 0
angle \sim N$$

o Rescaling the matrices as, $X^a \to \sqrt{N}X^a$ we see that

$$\mathcal{H}_{aff} = N \operatorname{Tr}\left(\sum_{a=1}^{6} \frac{1}{2} (D_t X^a)^2 + \frac{1}{2} (X^a)^2 + \sum_{a,b=1}^{6} \frac{g_{YM}^2 N}{8\pi^2} [X^a, X^b] [X^a, X^b]\right)$$

Will cost large energy in large N limit

1. Therefore we can now consider a reduced model of commuting matrix quantum mechanics:

$$S = \int dt \, \operatorname{Tr}\left[(D_t X^a)^2 - \frac{1}{2} (X^a)^2 \right] \qquad [X^a, X^b] = 0$$

2. To write Hamiltonian we need to take into account measure change in path integral:

$$H = \sum_{i} \left(-\frac{1}{2\mu^2} \nabla_i \mu^2 \nabla_i + \frac{1}{2} |\vec{x}_i|^2 \right) \qquad \mu^2 = \prod_{i < j} |\vec{x}_i - \vec{x}_j|^2$$
$$\vec{x}_i = (X_{ii}^1, \dots, X_{ii}^6)$$

1. Since the inner product contains the measure,

$$\langle \psi | \psi
angle = \int \prod_i d^6 x_i \mu^2 \psi^* \psi$$

We can re-scale the wavefunction as $\psi \rightarrow \psi/\mu$ and write

$$\begin{split} H &= \sum_{i} \left(-\frac{1}{2} \nabla_{i}^{2} + \frac{1}{2} |\vec{x}_{i}|^{2} \right) + V_{\text{eff}} \\ V_{\text{eff}} &= -6 \sum_{i \neq j} \frac{1}{|\vec{x}_{i} - \vec{x}_{j}|^{2}} + \sum_{i} \sum_{j,k \neq i} \frac{(\vec{x}_{i} - \vec{x}_{j}) \cdot (\vec{x}_{i} - \vec{x}_{k})}{|\vec{x}_{i} - \vec{x}_{j}|^{2} |\vec{x}_{i} - \vec{x}_{k}|^{2}} \end{split}$$

2. Ground state exactly known:

$$\psi_0 \sim \mu \exp\left(-rac{1}{2}\sum_i |\vec{x}_i|^2
ight)$$

1. Geometrical meaning of ground state (in large N limit)

$$\begin{split} |\psi_0|^2 &= \exp\left[-\sum_i |\vec{x}_i|^2 + \frac{1}{2} \sum_{i \neq j} \log |\vec{x}_i - \vec{x}_j|^2\right] \\ \text{arge N limit} &\to \exp\left[-\int d^6 x \rho(\vec{x}) |\vec{x}|^2 + \frac{1}{2} \int d^6 x d^6 y \rho(\vec{x}) \rho(\vec{y}) \log |\vec{x} - \vec{y}|^2\right] \end{split}$$

2. Can show that saddle point approximation on the inner product gives an $S^5 \subset R^6$ (hep-th/0507203, hep-th/0509015)

$$ho_0 = N rac{\delta(ert ec x ert - r_0)}{r_0^{2d-1} ext{Vol}(S^{2d-1})} \ r_0 = \sqrt{rac{N}{2}}$$

- 1. The claim is that this model should also describe graviton scattering in the bulk. (Note that this is a non-trivial interacting N- body system)
- 2. The proposal is that one should take an initial state similar to the one we studied at weak coupling:

$$\psi \sim e^{\sum_i \Omega_1(\vec{x}_i) + \sum_j \Omega_2(\vec{x}_j)} \psi_0$$

where each wavefunction should be approximately BPS in the large N limit

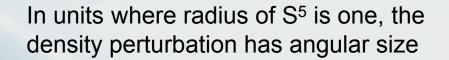
 $\langle H - J \rangle \approx 0$

 Berenstein, Cotta and Leonardi (hep-th/0605220, hepth/0702090, 0801.2739) have studied holomorphic states with

$$\Omega \sim \sum_k t_k z^k$$

and showed numerically that they are indeed approximately BPS.

2. To study high energy gravitons in flat space, one should use the moments t_k calculated before



$$\Delta\phi\sim 1/R\sim N^{-1/4}$$

- 1. Studying such scattering processes will require new numerical techniques (difficult to evolve in time).
- 2. We can estimate how important are the off-diagonal excitations: (Berenstein, Correa, Vazquez hep-th/0509015)

$$H_{ ext{off diag.}} \sim \sum_{i
eq j} \sum_{lpha} w_{i,j}^{(lpha)} (A_{lpha}^{\dagger})_i^j (A_{lpha})_j^i$$

$$w_{i,j}^{(\alpha)} = \sqrt{m_{lpha}^2 + rac{g_{YM}^2}{2\pi^2}} |ec{x}_i - ec{x}_j|^2$$

From conformal curvature coupling in action

 From commutator in action

 For two density perturbations separated a distance b in a flat space patch within the S⁵, the energy of the off-diagonal modes connecting them is

$$E_{\text{off-diag}} > \sqrt{g}b = l_s^{-2}b$$

2. To be able to ignore such modes one needs $E_{off-diag} << E_{cm}$ $b > l_s^2 E$

This is the well known bound to create long strings.

3. Not good enough for Black Hole formation. Need (Giddings, Eardley)

 $b \sim E^{1/7} \ll l_s^2 E$

- However, recently Giddings, Gross and Maharanara (0705.181) have shown that long strings are not important for black hole formation in high energy scattering.
- 2. Therefore, it is still possible that the reduced matrix model can describe such process.

Outlook

1. What have we learned?

- 1. Using the LLM dictionary, we learned to construct an initial state in SYM theory that is dual to a **regularized** Aichelburg-Sexl metric in flat space.
- 2. We learned how to put two such shock waves in a head on collision so that they will produce a classical black hole.
- 3. We made sure that the resulting trapped surface did not had any high curvatures.
- 4. I gave a proposal on how to study such scattering process at strong coupling in terms of a reduced matrix model.

2. What's next?

- 1. Understand how to compute scattering processes using the reduced matrix model (hard! But easier than full SYM...)
- 2. Understand the role of the off-diagonal modes in the black hole formation